

# SOLUTIONS TO ASSIGNMENTS 03

## 1. TENSOR ANALYSIS I: TENSOR ALGEBRA

- (a) Under a coordinate transformation  $x^\mu \rightarrow x'^\alpha = y^\alpha(x)$  a scalar (function)  $f(x)$  transforms as  $f \rightarrow f'$  with  $f'(y(x)) = f(x)$ , while according to the chain rule the partial derivatives transform with the Jacobian as

$$\partial_\alpha \equiv \partial_{y^\alpha} = J_\alpha^\mu \partial_\mu . \quad (1)$$

Thus

$$\partial_\alpha f'(y) = J_\alpha^\mu \partial_\mu f(x) \quad (2)$$

transforms as a covector.

- (b) By definition of a tensor we have (writing  $A_{\alpha\beta}$  instead of  $A'_{\alpha\beta}$  for simplicity, and suppressing the argument)

$$A_{\alpha\beta} = J_\alpha^\mu J_\beta^\nu A_{\mu\nu} \quad , \quad B^\beta = J_\rho^\beta B^\rho \quad , \quad (3)$$

and therefore

$$A_{\alpha\beta} B^\beta = J_\alpha^\mu J_\beta^\nu J_\rho^\beta A_{\mu\nu} B^\rho = J_\alpha^\mu \delta_\rho^\nu A_{\mu\nu} B^\rho = J_\alpha^\mu A_{\mu\nu} B^\nu \quad (4)$$

is a covector. The same kind of argument now shows that

$$A_{\alpha\beta} B^\alpha B^\beta = J_\rho^\alpha B^\rho J_\alpha^\mu A_{\mu\nu} B^\nu = \delta_\rho^\mu B^\rho A_{\mu\nu} B^\nu = B^\mu A_{\mu\nu} B^\nu \quad (5)$$

is a scalar.

- (c) The invariance of  $V(x)$  under coordinate transformations follows from the fact that partial derivatives are covectors and that they are contracted with a vector to form the field  $V(x)$ ,

$$V^\alpha \partial_\alpha = J_\mu^\alpha J_\alpha^\nu V^\mu \partial_\nu = \delta_\mu^\nu V^\mu \partial_\nu = V^\mu \partial_\mu . \quad (6)$$

Likewise for a covector:

$$dy^\alpha = J_\nu^\alpha dx^\nu \quad \Rightarrow \quad A_\alpha dy^\alpha = J_\alpha^\mu J_\nu^\alpha A_\mu dx^\nu = A_\mu dx^\mu . \quad (7)$$

## 2. THE EFFECTIVE GEODESIC POTENTIAL

Starting with the metric

$$ds^2 = -f(r)dt^2 + f(r)^{-1}dr^2 + r^2d\Omega^2 \quad , \quad f(r) = 1 + 2\phi(r) \quad (8)$$

one implements the following steps:

- the Lagrangian  $\mathcal{L}$  is conserved,

$$-f(r)\dot{t}^2 + f(r)^{-1}\dot{r}^2 + r^2(\dot{\theta}^2 + \sin^2\theta\dot{\phi}^2) = \epsilon \quad (9)$$

where  $\epsilon = -1, 0$  for massive (massless) particles.

- by spherical symmetry, angular momentum is conserved, thus the motion is planar, and one can choose the coordinates such that this motion takes place in the equatorial plane  $\theta = \pi/2$ ,  $\dot{\theta} = 0$ , leading to

$$-f(r)\dot{t}^2 + f(r)^{-1}\dot{r}^2 + r^2\dot{\phi}^2 = \epsilon \quad (10)$$

- rotational and time-translational symmetry lead to the conserved quantities

$$E = f(r)\dot{t} \quad L = r^2\dot{\phi} \quad (11)$$

(energy and angular momentum), and using these equations to eliminate  $\dot{t}$  and  $\dot{\phi}$  from the Lagrangian, one finds

$$-E^2 f(r)^{-1} + f(r)^{-1}\dot{r}^2 + L^2/r^2 = \epsilon \quad (12)$$

Multiplying by  $f(r)$  and rearranging, this gives

$$\dot{r}^2 + f(r)L^2/r^2 - \epsilon f(r) = E^2 \quad (13)$$

- This already has the desired form of an effective Newtonian potential equation, but it is typically more useful to separate the constant (asymptotically Minkowski) part of  $f(r)$  from the rest. Thus, with  $f(r) = 1 + 2\phi(r)$  one has

$$\frac{1}{2}\dot{r}^2 + V_{eff}(r) = E_{eff} \quad (14)$$

where

$$V_{eff}(r) \equiv V(r) + L^2/2r^2 = \phi(r)(-\epsilon + L^2/r^2) + L^2/2r^2 \quad (15)$$

and

$$E_{eff} = (E^2 + \epsilon)/2 \quad (16)$$