

1. ORBIT EQUATION FOR GENERAL STATIC SPHERICALLY SYMMETRIC METRICS

First we extract the constants of motion from the Lagrangian :

$$\begin{aligned} \mathcal{L} = \frac{1}{2}g_{\mu\nu}\dot{x}^\mu\dot{x}^\nu = \frac{\epsilon}{2} &= -\frac{1}{2}A(r)\dot{t}^2 + \frac{1}{2}B(r)\dot{r}^2 + \frac{1}{2}r^2\dot{\theta}^2 + \frac{1}{2}r^2\sin(\theta)^2\dot{\phi}^2 \\ &\Rightarrow -\frac{\partial\mathcal{L}}{\partial\dot{t}} = E = A(r)\dot{t} \end{aligned} \quad (1)$$

$$\Rightarrow \frac{\partial\mathcal{L}}{\partial\dot{\phi}} = L = r^2\sin(\theta)^2\dot{\phi} \quad (2)$$

$$\text{and we fix } \theta = \theta_0 = \frac{\pi}{2} \Rightarrow \mathcal{L} = -\frac{E^2}{2A(r)} + \frac{B(r)}{2}\dot{r}^2 + \frac{L^2}{2r^2} = \frac{\epsilon}{2} \quad (3)$$

We have $r = r(\tau)$ but we want to express r as a function of ϕ , as we do so this implies that $\dot{r} = r'\dot{\phi} = \frac{Lr'}{r^2}$ and we get with $r = r(\phi)$ and $r' = \frac{dr}{d\phi}$:

$$-\frac{E^2}{L^2A(r)} + B(r)\frac{r'^2}{r^4} + \frac{1}{r^2} = \frac{\epsilon}{L^2} \quad (4)$$

Finally, we change variable from r to $u = 1/r$ ($\Rightarrow r' = -\frac{u'}{u^2}$) and (4) becomes :

$$\tilde{B}(u)u'^2 + u^2 = \frac{\epsilon}{L^2} + \frac{E^2}{L^2\tilde{A}(u)} \quad (5)$$

where $\tilde{B}(u) = B(r(u)) = B(\frac{1}{u})$ and the same for $\tilde{A}(u)$.

2. RADIAL FALL & THE REPULSIVE REISSNER-NORDSTRÖM CORE

We use equation (3) with $\epsilon = -1$, $A = B^{-1} = 1 - \frac{2m}{r} + \frac{q^2}{r^2}$ and $L = 0$ which after multiplying both sides by A gives :

$$-\frac{E^2}{2} + AB\frac{\dot{r}^2}{2} = -\frac{A}{2} \quad (6)$$

$$\frac{\dot{r}^2}{2} = \frac{E^2}{2} - \frac{A}{2} \quad (7)$$

$$\frac{\dot{r}^2}{2} - \frac{m}{r} + \frac{q^2}{2r^2} = \frac{E^2}{2} - \frac{1}{2} \quad (8)$$

$$\frac{\dot{r}^2}{2} + V_{eff}(r) = \mathcal{E} \quad (9)$$

We observe that for $q \neq 0$ the dominant term in the effective potential for a small radius is now positive and we have $V_{eff}(r \rightarrow 0) \rightarrow \infty$ such that infalling particles cannot reach $r = 0$ (as they would need an infinite amount of energy).

3. TENSOR ANALYSIS IV: THE RIEMANN CURVATURE TENSOR

(a) On a scalar we have $\nabla_\mu \phi = \partial_\mu \phi$, therefore :

$$[\nabla_\mu, \nabla_\nu] \phi = (\nabla_\mu \nabla_\nu - \nabla_\nu \nabla_\mu) \phi \quad (10)$$

$$= (\nabla_\mu \partial_\nu - \nabla_\nu \partial_\mu) \phi \quad (11)$$

$$= (\partial_\mu \partial_\nu - \partial_\nu \partial_\mu) \phi - (\Gamma_{\mu\nu}^\rho - \Gamma_{\nu\mu}^\rho) \partial_\rho \phi \quad (12)$$

$$= 0 \quad (13)$$

(b) We just have to compute :

$$\begin{aligned} [\nabla_\mu, \nabla_\nu] V^\lambda &= \nabla_\mu [\partial_\nu V^\lambda + \Gamma_{\nu\rho}^\lambda V^\rho] - (\mu \leftrightarrow \nu) \\ &= \partial_\mu \partial_\nu V^\lambda + \partial_\mu \Gamma_{\nu\rho}^\lambda V^\rho + \Gamma_{\mu\rho}^\lambda \partial_\nu V^\rho - \Gamma_{\mu\nu}^\rho \partial_\rho V^\lambda + \Gamma_{\mu\sigma}^\lambda \Gamma_{\nu\rho}^\sigma V^\rho - \Gamma_{\mu\nu}^\sigma \Gamma_{\sigma\rho}^\lambda V^\rho - (\mu \leftrightarrow \nu) \\ &= \partial_\mu \Gamma_{\nu\rho}^\lambda V^\rho + \Gamma_{\mu\rho}^\lambda \partial_\nu V^\rho + \Gamma_{\mu\sigma}^\lambda \Gamma_{\nu\rho}^\sigma V^\rho - (\mu \leftrightarrow \nu) \\ &= (\partial_\mu \Gamma_{\nu\rho}^\lambda) V^\rho + \Gamma_{\nu\rho}^\lambda \partial_\mu V^\rho + \Gamma_{\mu\rho}^\lambda \partial_\nu V^\rho + \Gamma_{\mu\sigma}^\lambda \Gamma_{\nu\rho}^\sigma V^\rho - (\mu \leftrightarrow \nu) \\ &= (\partial_\mu \Gamma_{\nu\rho}^\lambda) V^\rho + \Gamma_{\mu\sigma}^\lambda \Gamma_{\nu\rho}^\sigma V^\rho - (\mu \leftrightarrow \nu) \\ &= [\partial_\mu \Gamma_{\nu\rho}^\lambda - \partial_\nu \Gamma_{\mu\rho}^\lambda + \Gamma_{\mu\sigma}^\lambda \Gamma_{\nu\rho}^\sigma - \Gamma_{\nu\sigma}^\lambda \Gamma_{\mu\rho}^\sigma] V^\rho = R_{\rho\mu\nu}^\lambda V^\rho \end{aligned} \quad (14)$$

where in the third equality we dropped all the (μ, ν) -symmetric terms killed by the subtraction with the indices exchanged.